

AIRS Calibration Update and Radiometric Uncertainty Estimate

Thomas S. Pagano^a

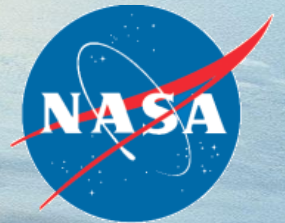
Evan M. Manning^a, Steven E. Broberg^a, Hartmut Aumann^a,
Margie Weiler^b, Kenneth Overoye^b

^aJet Propulsion Laboratory, California Institute of Technology,
4800 Oak Grove Dr., Pasadena, CA 91109
tpagano@jpl.nasa.gov, (818) 393-3917, <http://airs.jpl.nasa.gov>

^bBAE Systems, Nashua, New Hampshire, 03064 (Retired)

GSICS Annual Meeting 2019, Frascati, Italy

March 7, 2019





Agenda

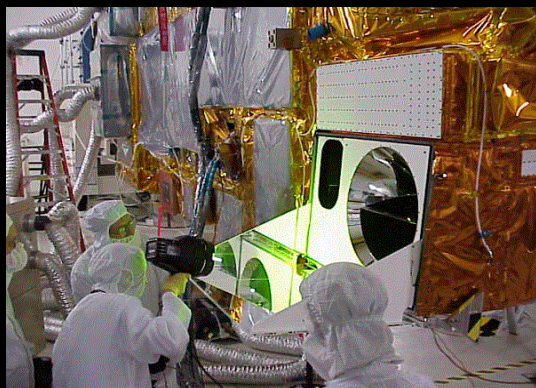


- AIRS on Aqua Healthy. Expect mission to continue beyond 2022.
- Continuity from operational weather sounders: CrIS, IASI
- AIRS designed and tested to produce SI traceable radiances
- Radiometric Calibration Coefficients Updated
 - Current operational version using pre-launch coefficients (V5)
 - Updates provided to polarization, emissivity, nonlinearity (V7k)
- V7k Improvements seen in:
 - Reduced Bias and L/R Assymetry in Cold Scenes
 - Agrees better with CrIS
 - Reduction in Radiometric Accuracy



The Aqua Spacecraft

Launched May 4, 2002



Moderate Resolution Imaging Spectroradiometer (MODIS)
GSFC/Raytheon



Atmospheric Infrared Sounder (AIRS)
JPL/BAE SYSTEMS



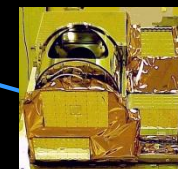
AQUA Spacecraft
GSFC/NGST



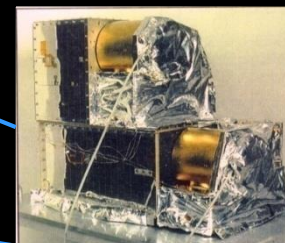
Advanced Microwave Scanning Radiometer (AMSR-E)
MSFC/JAXA



Advanced Microwave Sounding Units (AMSU-A/B)
JPL/Aerojet



Humidity Sounder from Brazil (HSB)
JPL/Aerojet



Clouds and Earth Radiant Energy System (CERES)
LaRC/NGST



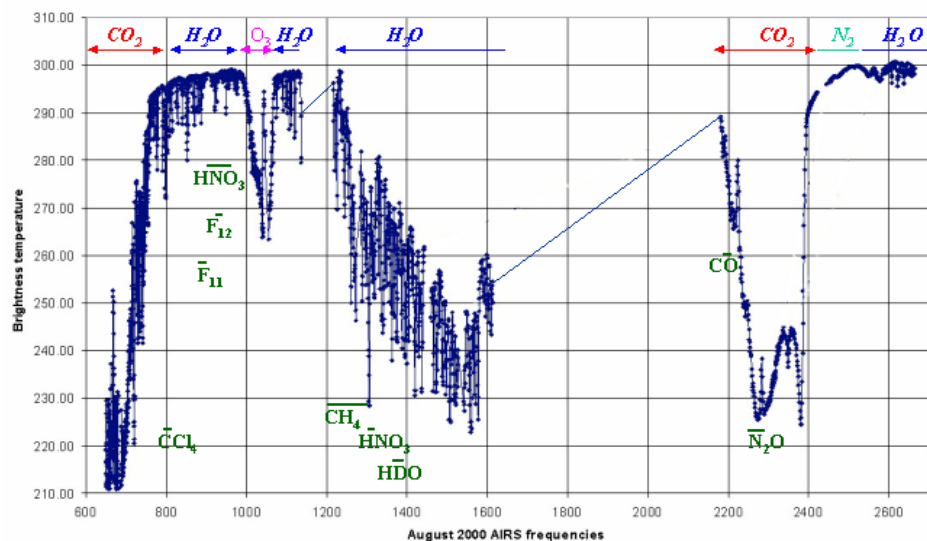


IR Sounders Support Weather Forecasting and Climate Science

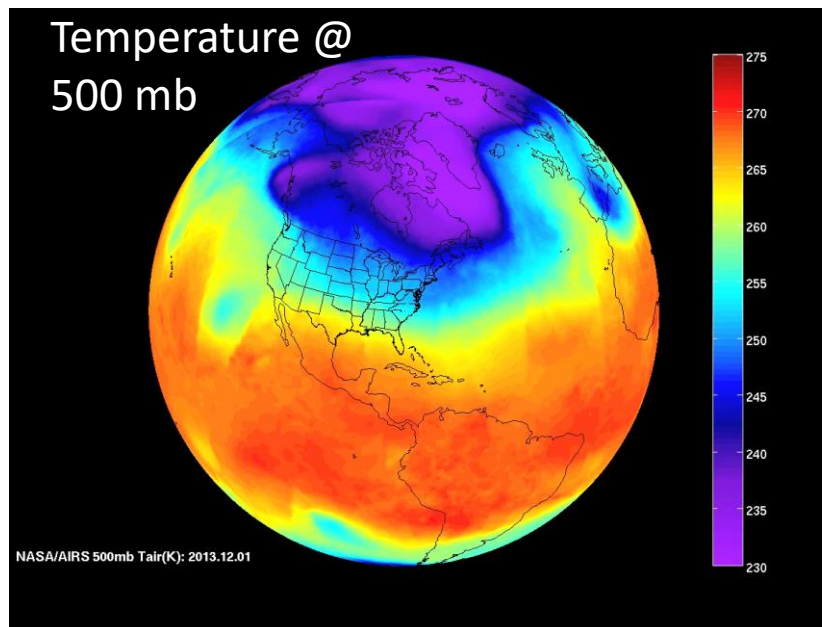


AIRS Channels for Tropical Atmosphere with $T_{\text{surf}} = 301\text{K}$

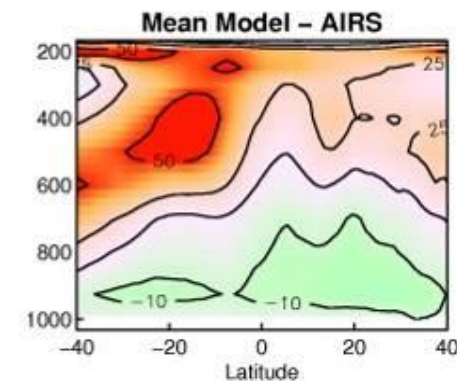
Full Spectrum



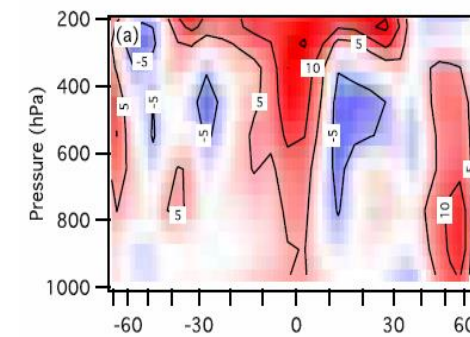
Temperature @
500 mb



Water Vapor Climatology
(Pierce, Scripps, 2006)



Water Vapor Feedback
(Dessler, Texas A&M, 2008)



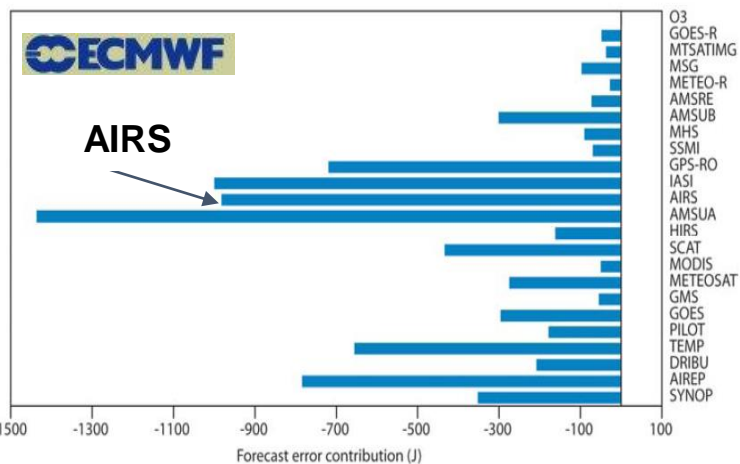
JPL/GSFC



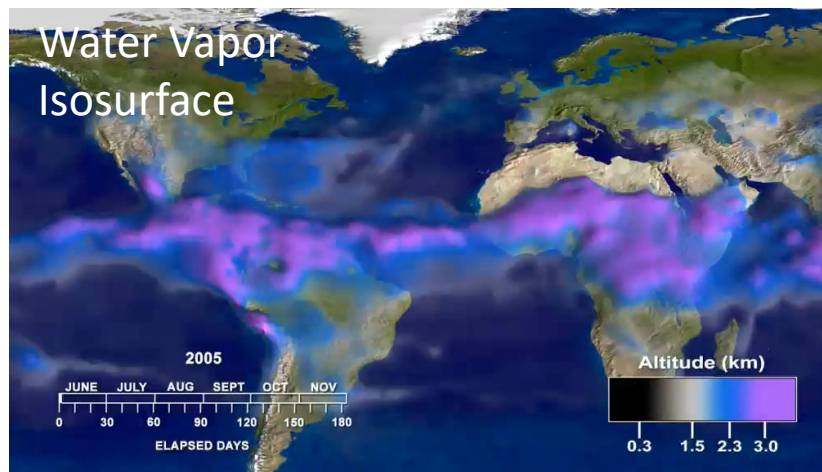
NOAA
NESDIS/NCEP



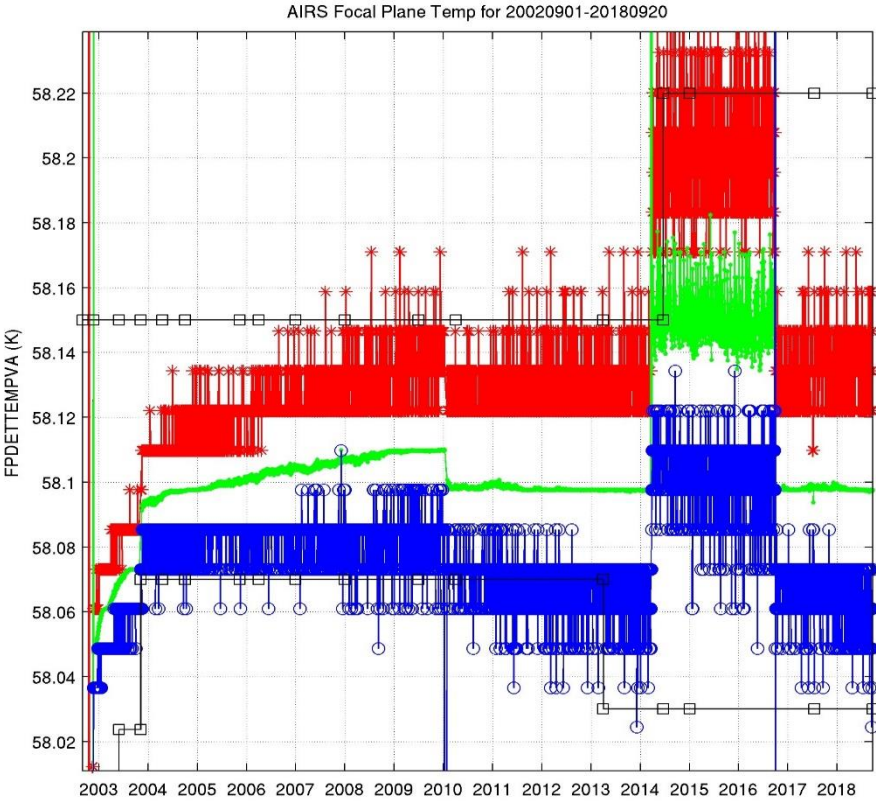
JCSDA



Water Vapor
Isosurface

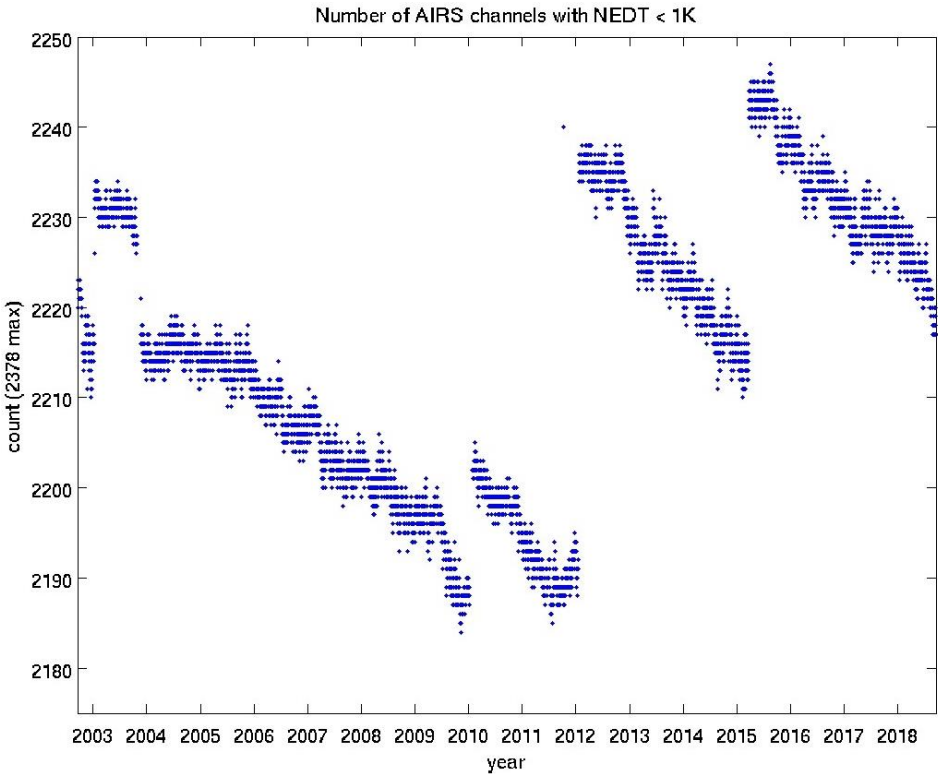


Minimal Trends in Instrument Telemetry



Focal Plane temperature has remained within a 0.25 K range throughout the AIRS mission. Note: Median temperature (green) rose ~0.06 K during 2014-2016 Cooler A anomaly

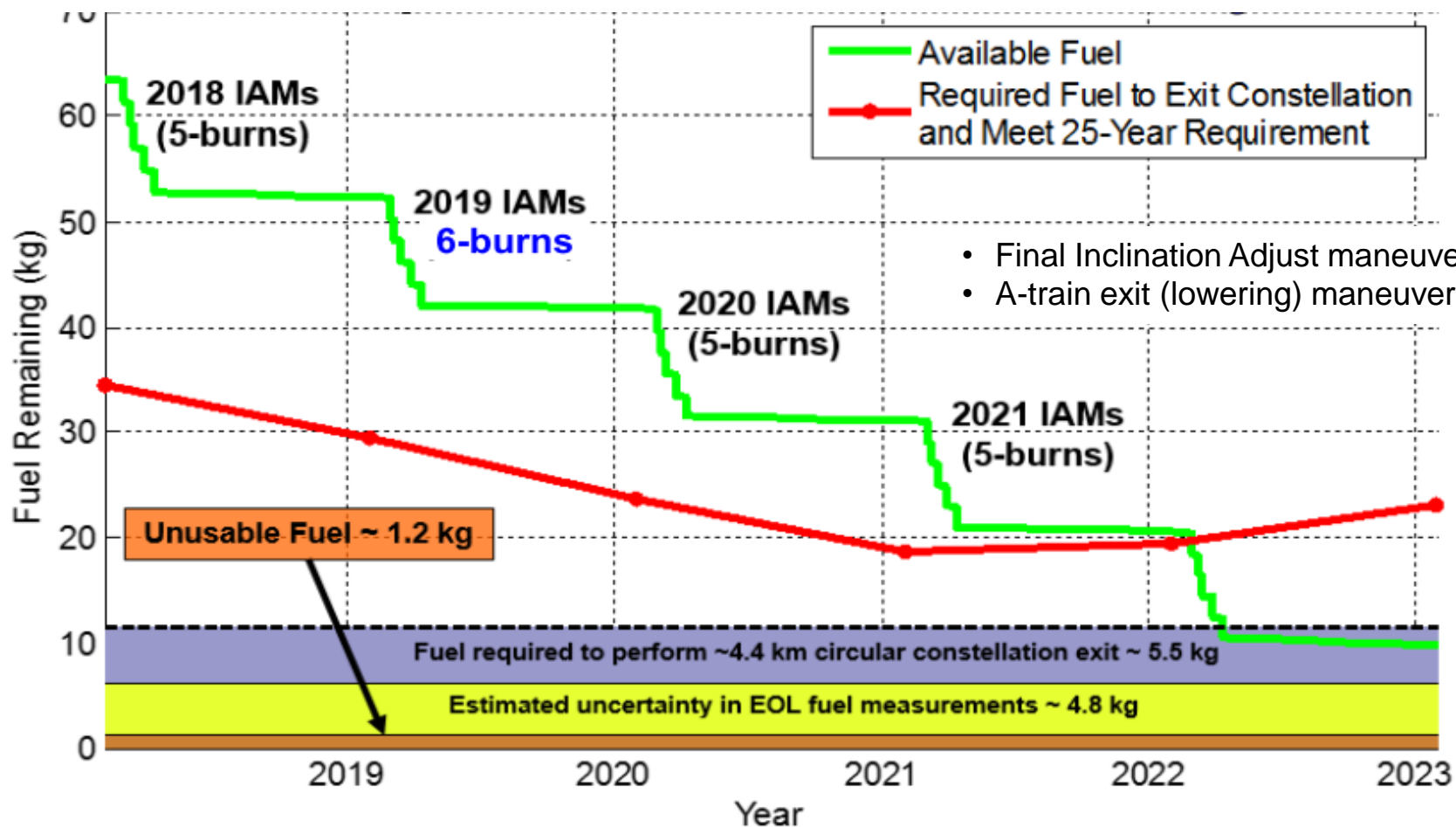
Detector Operability Maintained Throughout Mission



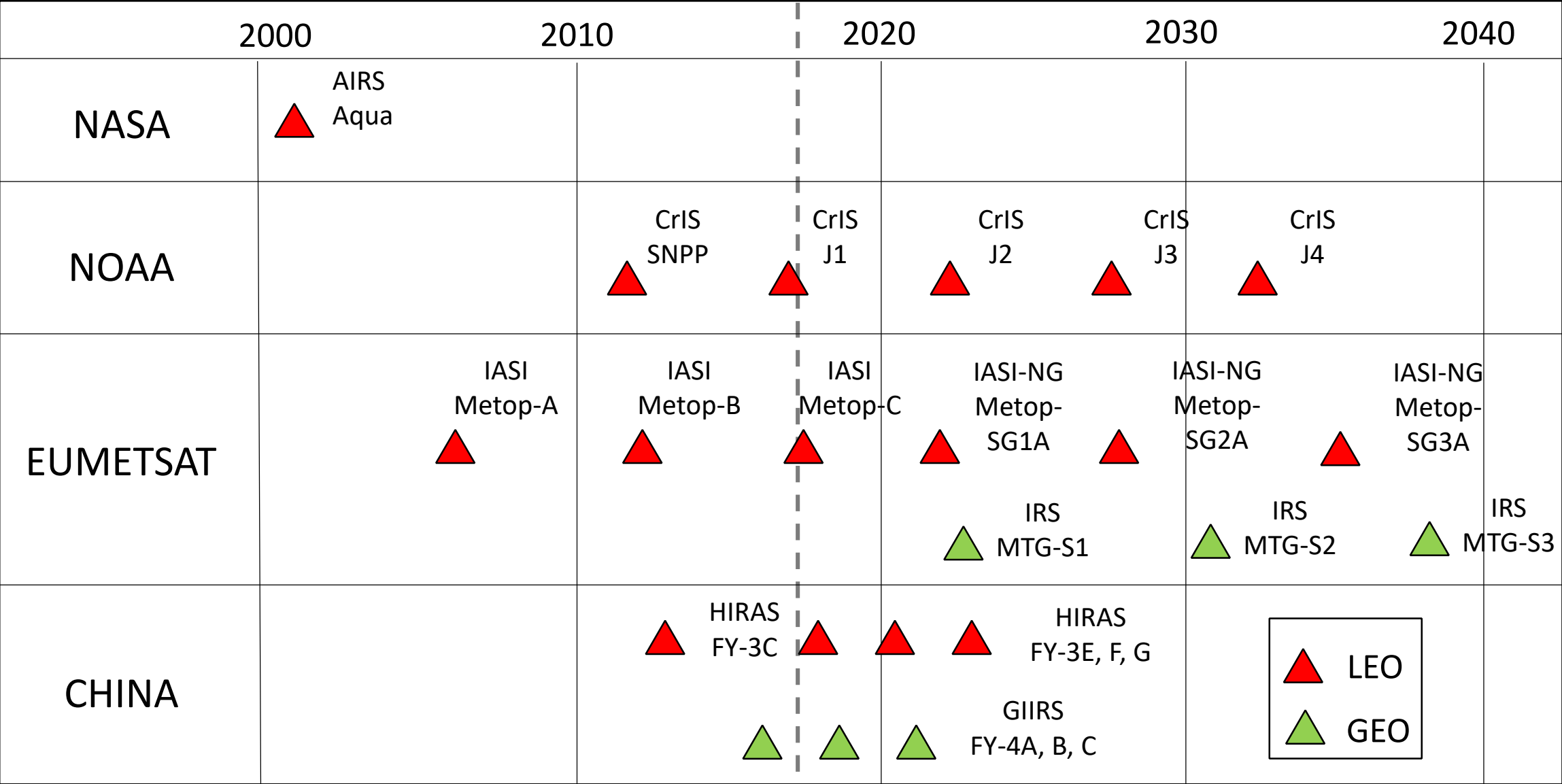
GainTable loads

2003-01-10	2013-06-10
2003-11-18	2015-03-23
2012-01-21	2018-xx-xx

- Aqua Flight Operations Team is continuing to test alternate orbit-lowering maneuver(s)
- Post-2022 orbit will also have thermal impact on AIRS – specifically, 2nd stage heater will need to draw more power to maintain spectrometer temperature
- AIRS is expected to last the life of the spacecraft



- Final Inclination Adjust maneuvers: March-April 2021
- A-train exit (lowering) maneuver: March 2022

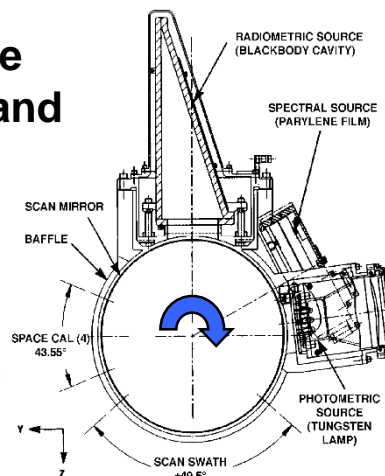


AIRS Features

- Orbit: 705 km, 1:30pm, Sun Synch
- Pupil Imaging IFOV : $1.1^\circ \times 0.6^\circ$
(13.5 km x 7.4 km)
- Scanner Rotates about Optical Axis
(Constant AOI on Mirror)
- Full Aperture OBC Blackbody, $\epsilon > 0.998$
- Full Aperture Space View
- Solid State Grating Spectrometer
- Temperature Controlled Spectrometer: 158K
- Actively Cooled FPAs: 60K
- No. Channels: 2378 IR, 4 Vis/NIR
- Mass: 177Kg,
Power: 256 Watts,
Life: 5 years (7 years goal)

Full Aperture Blackbody and Space View

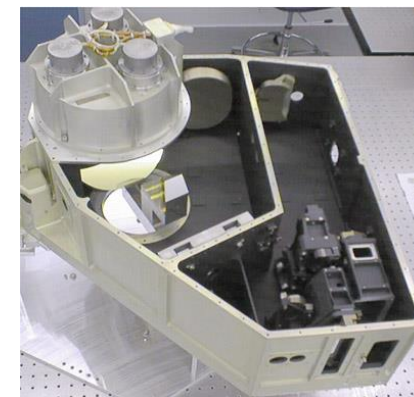
Isolated Scan Cavity



Temperature Controlled Instrument



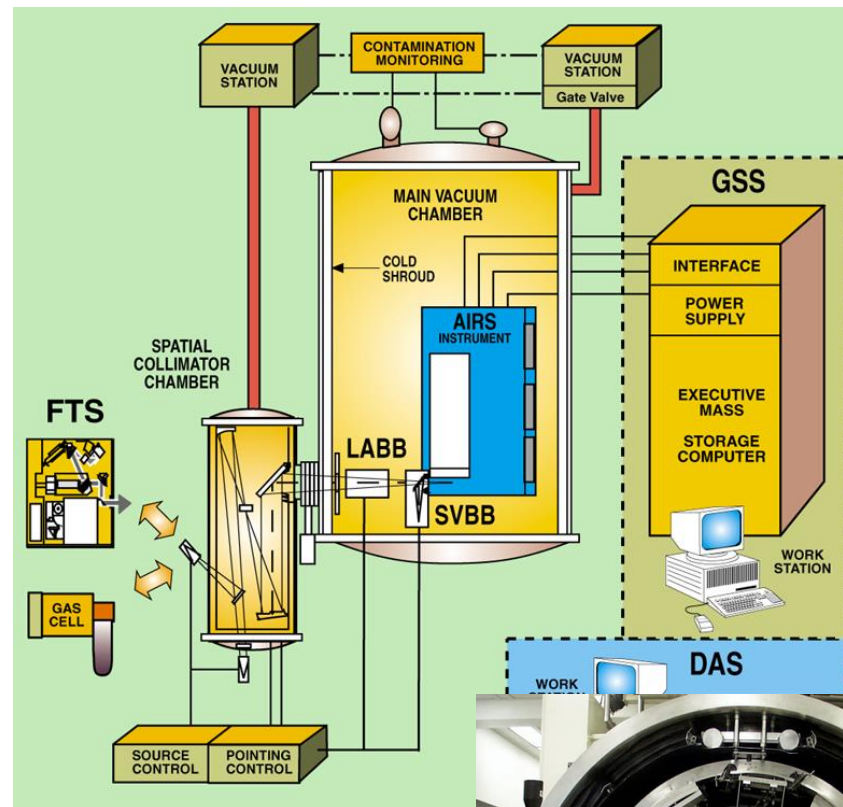
Active Detector Cooling



Grating Spectrometer

IR Spectral Range:
3.74-4.61 μm , 6.2-8.22 μm ,
8.8-15.4 μm
IR Spectral Resolution:
 $\approx 1200 (\lambda/\Delta\lambda)$
No. IR Channels: 2378 IR

- Radiometric Response
 - Emissivity, Nonlinearity
 - Stray Light, Polarization
 - Scan Angle Dependence in TVAC
 - Transfer to On-Board Blackbody
 - 2 TVAC Cycles
- Spectral Response
 - SRF Characterization with FTS
 - Channel Spectra Characterized
- Spatial Response
 - Top-hat Functions All Channels
 - Stray Light Excellent
 - Far Field Response Excellent
- Good Documentation
 - Over 400 Design File Memos



AIRS Radiometric Transfer Equations

Scene Radiance

$$L_{ev} = L_o(\theta) + \frac{c_0 + c'_1(dn_{ev} - dn_{sv}) + c_2(dn_{ev} - dn_{sv})^2}{[1 + p_r p_t \cos 2(\theta - \delta)]}$$

Mirror Polarization Contribution

$$L_o(\theta) = \frac{L_{sm} p_r p_t [\cos 2(\theta - \delta) - \cos 2(\theta_{sv,i} - \delta)]}{[1 + p_r p_t \cos 2(\theta - \delta)]}$$

Gain Term

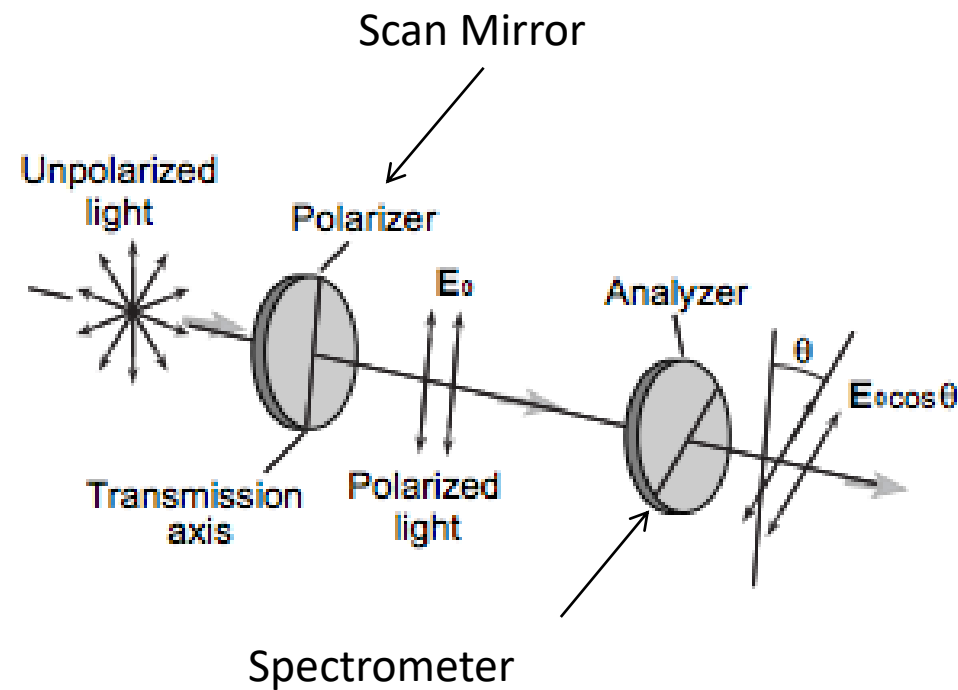
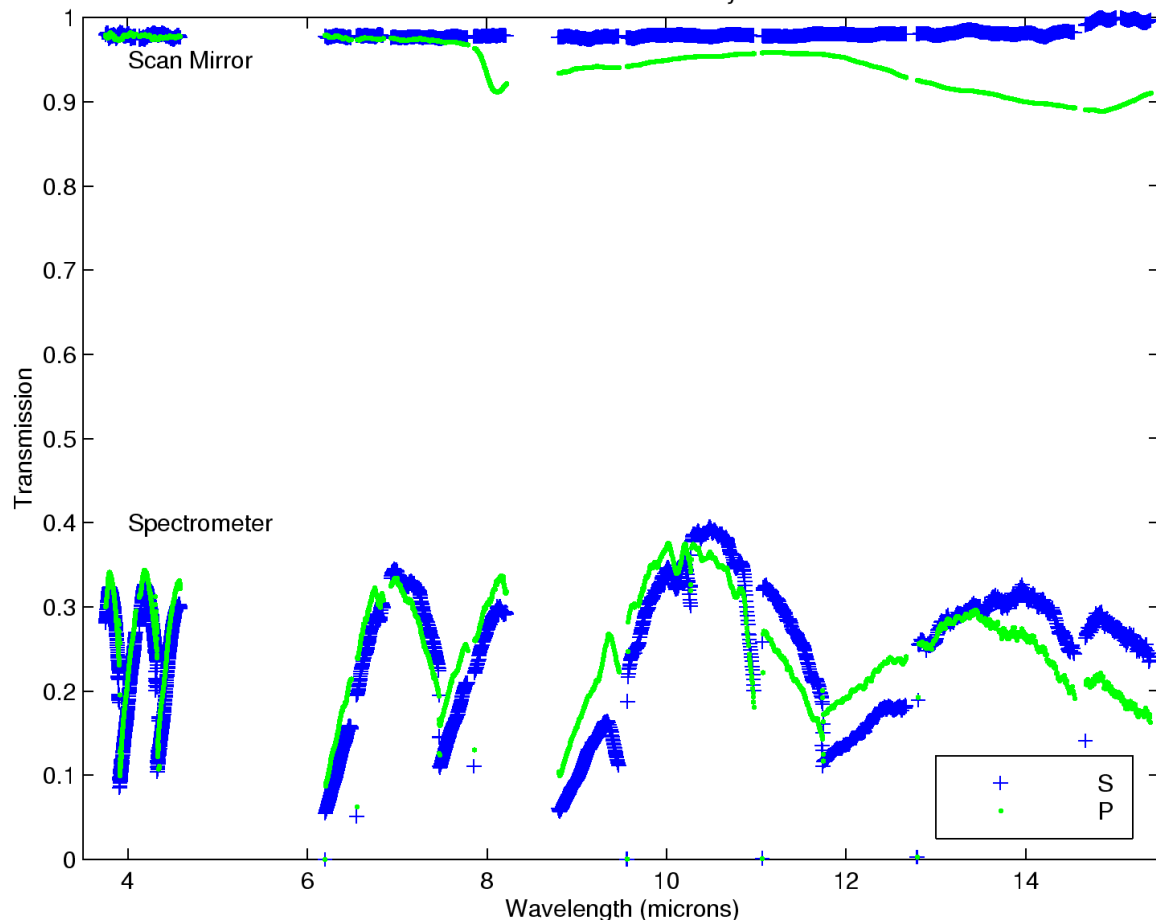
$$c'_1 = \frac{[\varepsilon_{obc} P_{obc} - L_o(\theta_{obc})][1 + p_r p_t \cos 2\delta] - c_2(dn_{obc} - dn_{sv})^2 - c_0}{(dn_{obc} - dn_{sv})}$$

L_{ev} = Spectral Radiance in the Earth Viewport (W/m²-sr-μm)
 L_{sm} = Spectral Radiance of the Scan Mirror for Unity Emissivity at T_{sm} (W/m²-sr-μm)
 c_0 = Instrument offset (W/m²-sr- μm)
 c_1 = Instrument gain (W/m²-sr- μm-counts)
 c_2 = Instrument nonlinearity (W/m²-sr- μm-counts²)
 dn_{ev} = Digital counts while viewing Earth for each footprint and scan (counts)
 dn_{sv} = Digital counts while viewing Space for each scan (counts)
 $p_r p_t$ = Product of scan mirror and spectrometer polarization diattenuation (unitless)
 θ = Scan Angle measured from nadir (radians)
 δ = Phase of spectrometer polarization (radians)
 P_{obc} = Plank Blackbody function of the OBC blackbody at temperature T_{obc} (W/m²-sr-μm)
 T_{obc} = Telemetered temperature of the OBC blackbody (K) with correction of +0.3K.
 ε_{obc} = Effective Emissivity of the blackbody
 dn_{obc} = Digital number signal from the AIRS while viewing the OBC Blackbody

V7k

- T. Pagano et al., "Reducing uncertainty in the AIRS radiometric calibration", Proc. SPIE 10764-23, San Diego, CA (2018)
- T. Pagano et al., "Pre-Launch and In-flight Radiometric Calibration of the Atmospheric Infrared Sounder (AIRS)," IEEE TGRS, Volume 41, No. 2, February 2003, p. 265
- T. Pagano, H. Aumann, K. Overoye, "Level 1B Products from the Atmospheric Infrared Sounder (AIRS) on the EOS Aqua Spacecraft", Proc. ITOVS, October 2003

AIRS Theoretical Transmissions Modified by Measured Polarization

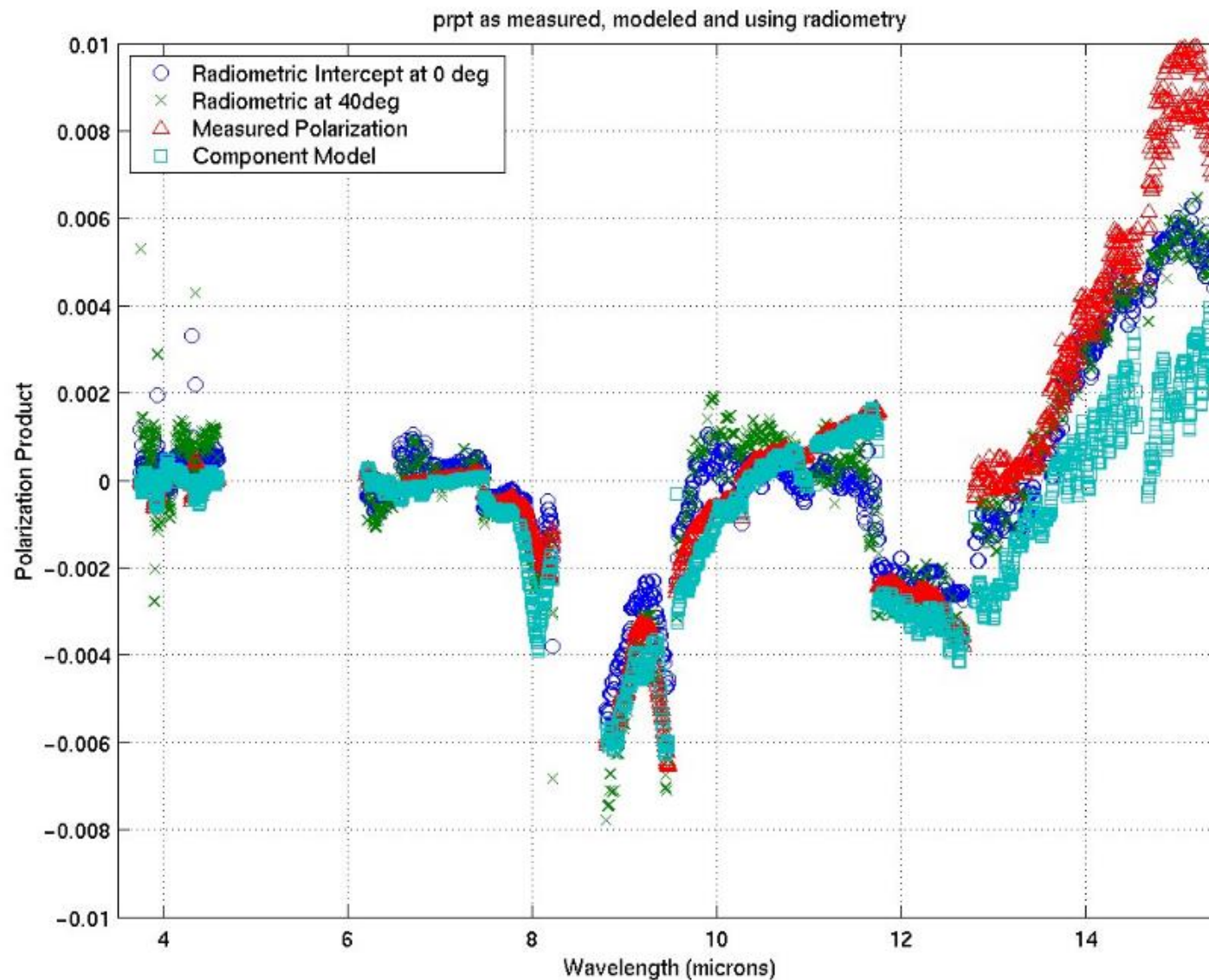


$$L_{ev} = L_o(\theta) + \frac{c_0 + c'_1(dn_{ev} - dn_{sv}) + c_2(dn_{ev} - dn_{sv})^2}{[1 + p_r p_t \cos 2(\theta - \delta)]}$$

Assumes all space views at 90°

$$L_o(\theta) = \frac{L_{sm} p_r p_t [\cos 2(\theta - \delta) + \cos 2\delta]}{[1 + p_r p_t \cos 2(\theta - \delta)]}$$

V5 Coefficients
Determined Pre-
Launch and Not
Changed Since



High Uncertainty
Especially in LWIR

V5 is average of
“Measured” and
“Modeled”
polarization, with
Phase = 0

New Method Uses Multiple Space Views to Derive Polarization

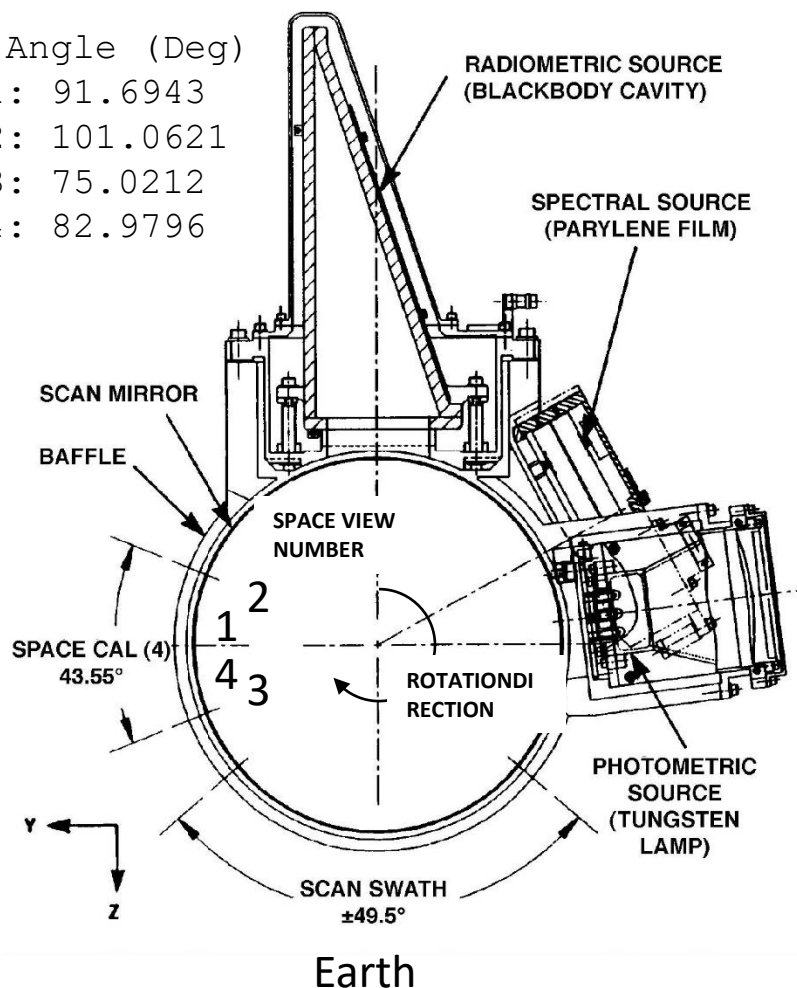
SV Angle (Deg)

SV1: 91.6943

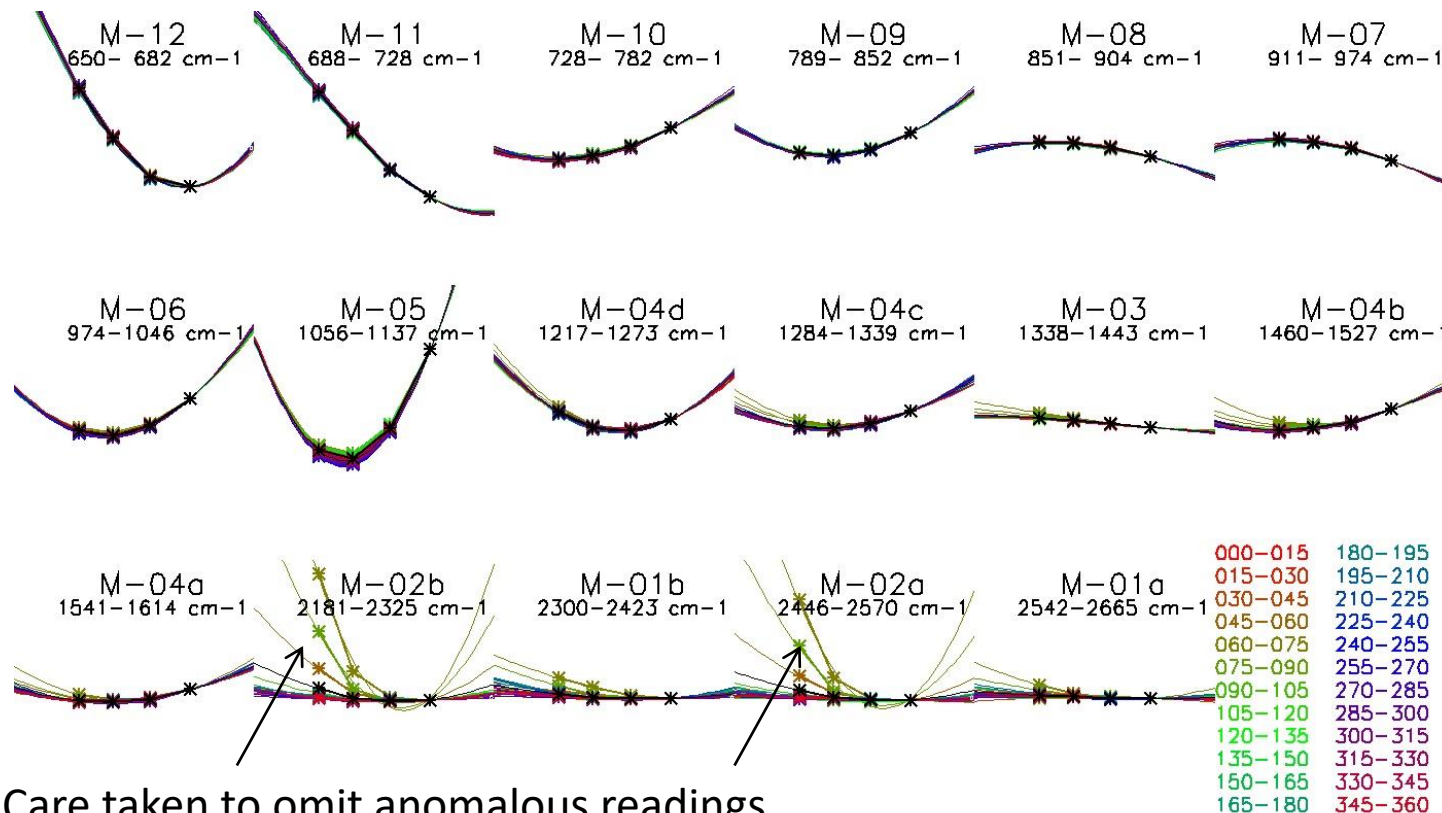
SV2: 101.0621

SV3: 75.0212

SV4: 82.9796



Every space view in the mission used to give 171 monthly averages

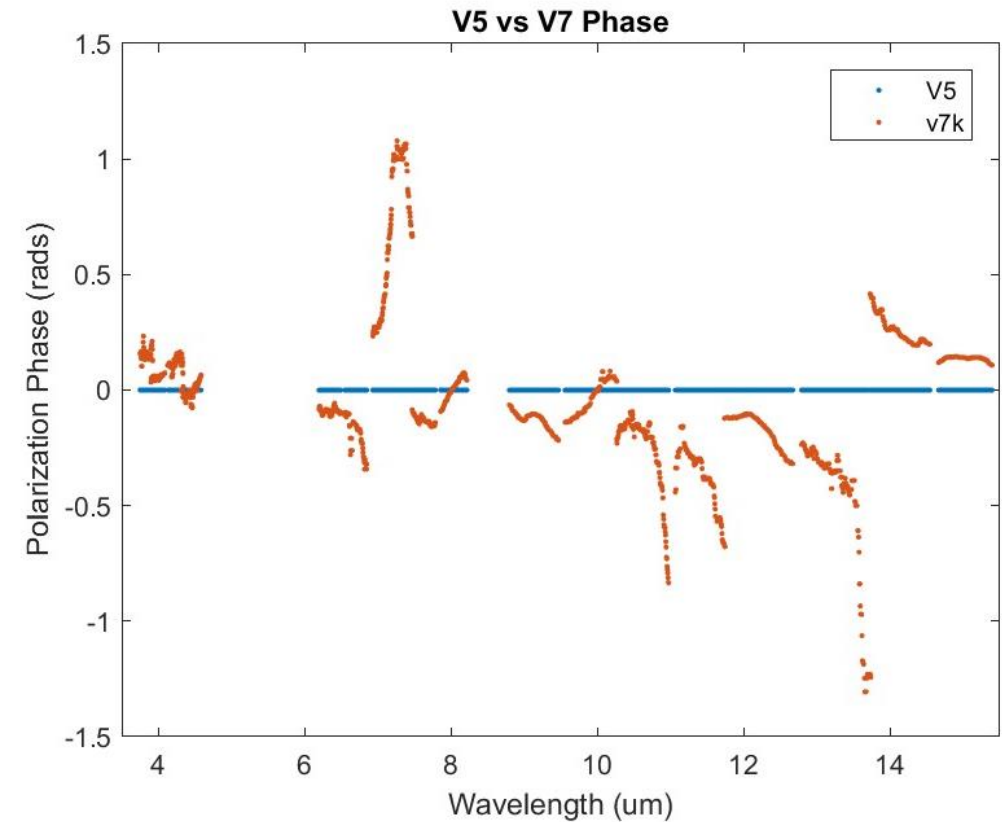
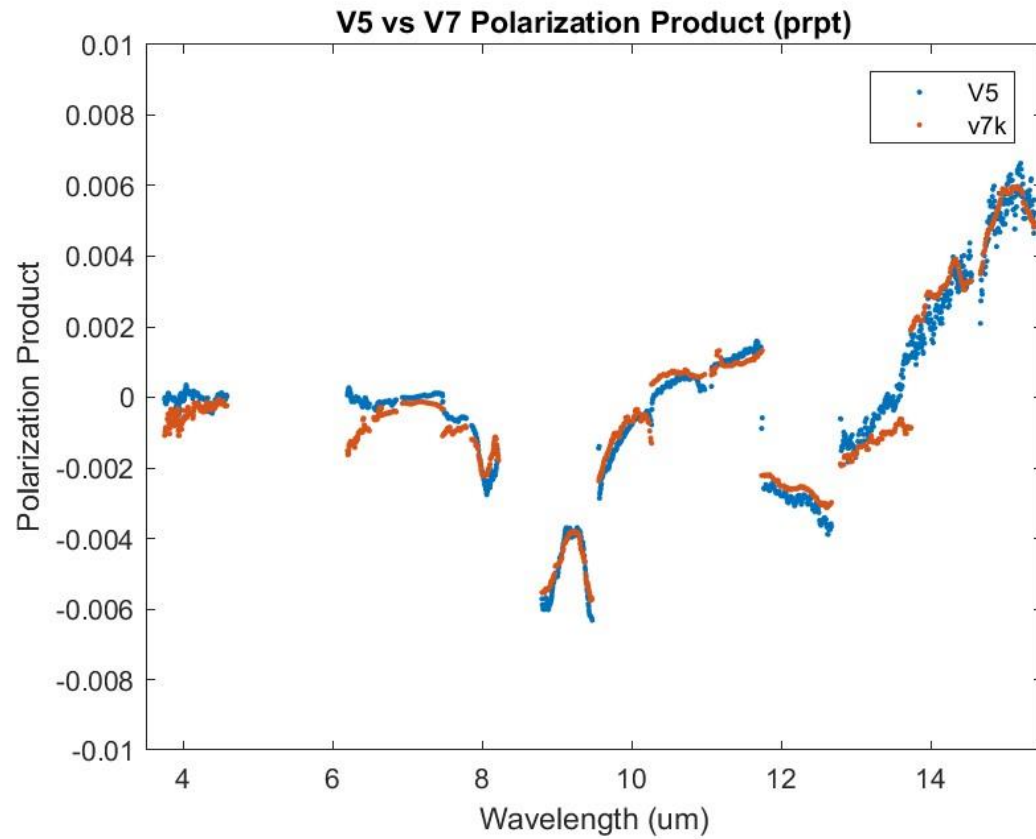


Care taken to omit anomalous readings

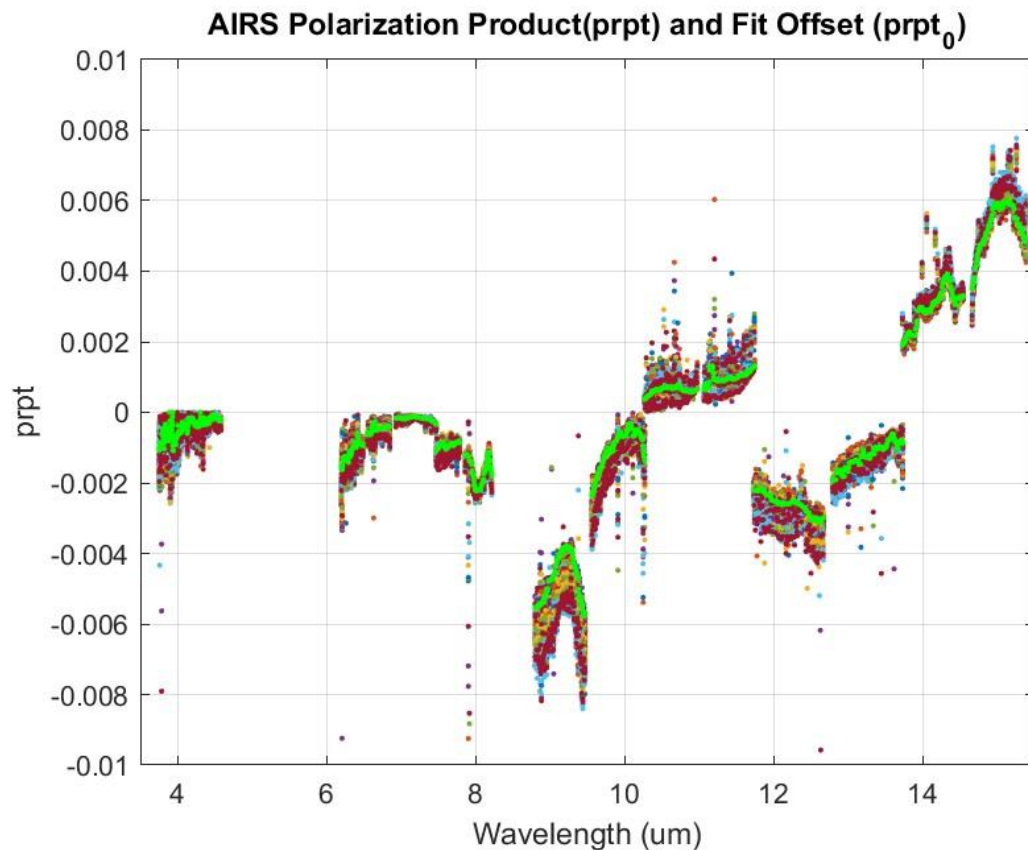
$$(dn_{sv,i} - dn_{sv,1})c_1' = -L_{sm}p_r p_t [\cos 2\theta_{sv,i} \cos 2\delta + \sin 2\theta_{sv,i} \sin 2\delta + \cos 2\delta]$$

$p_r p_t$

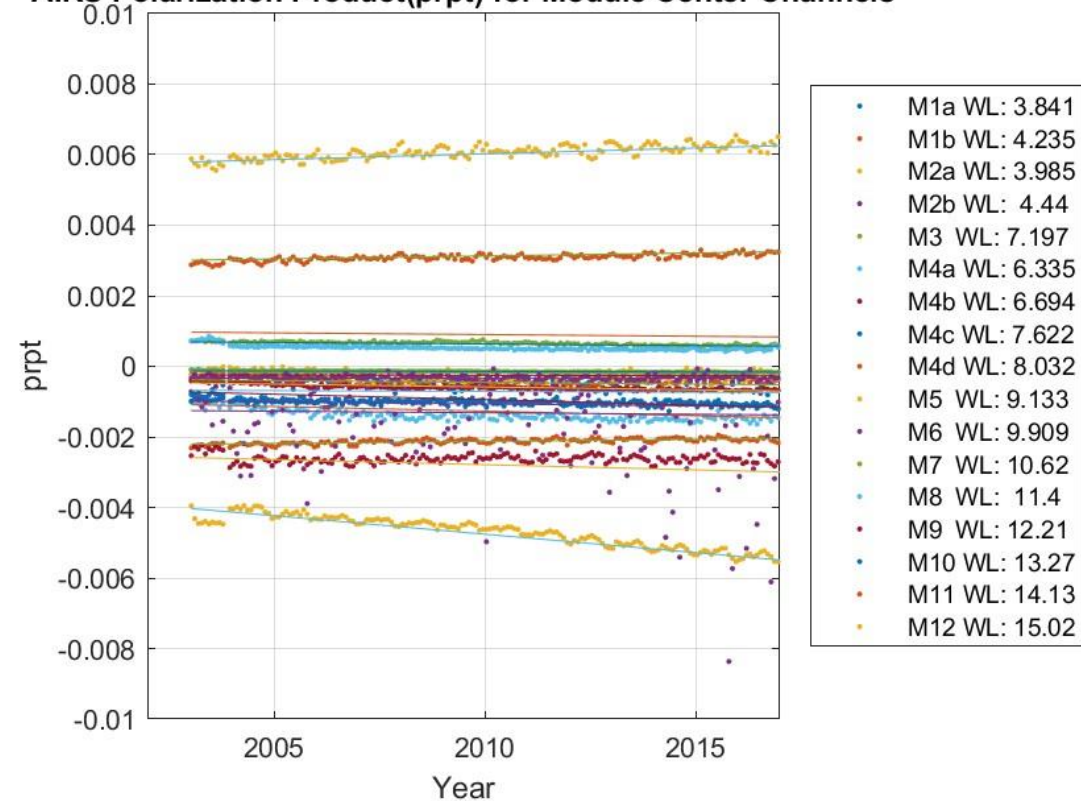
phase



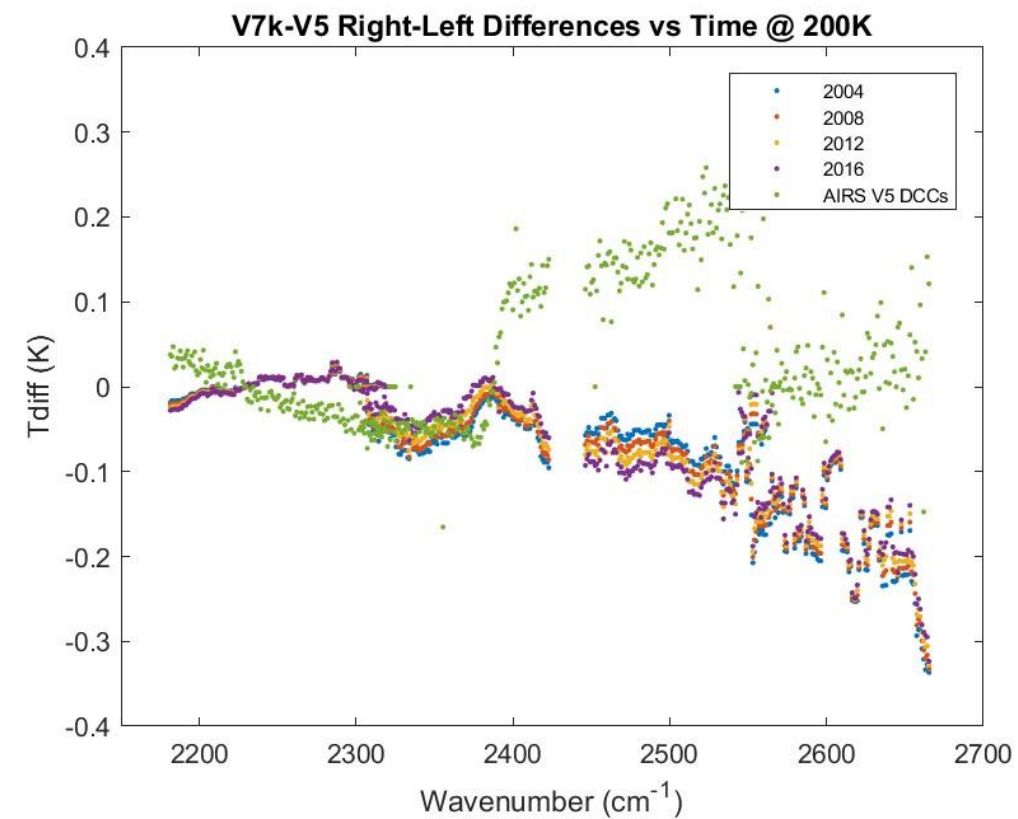
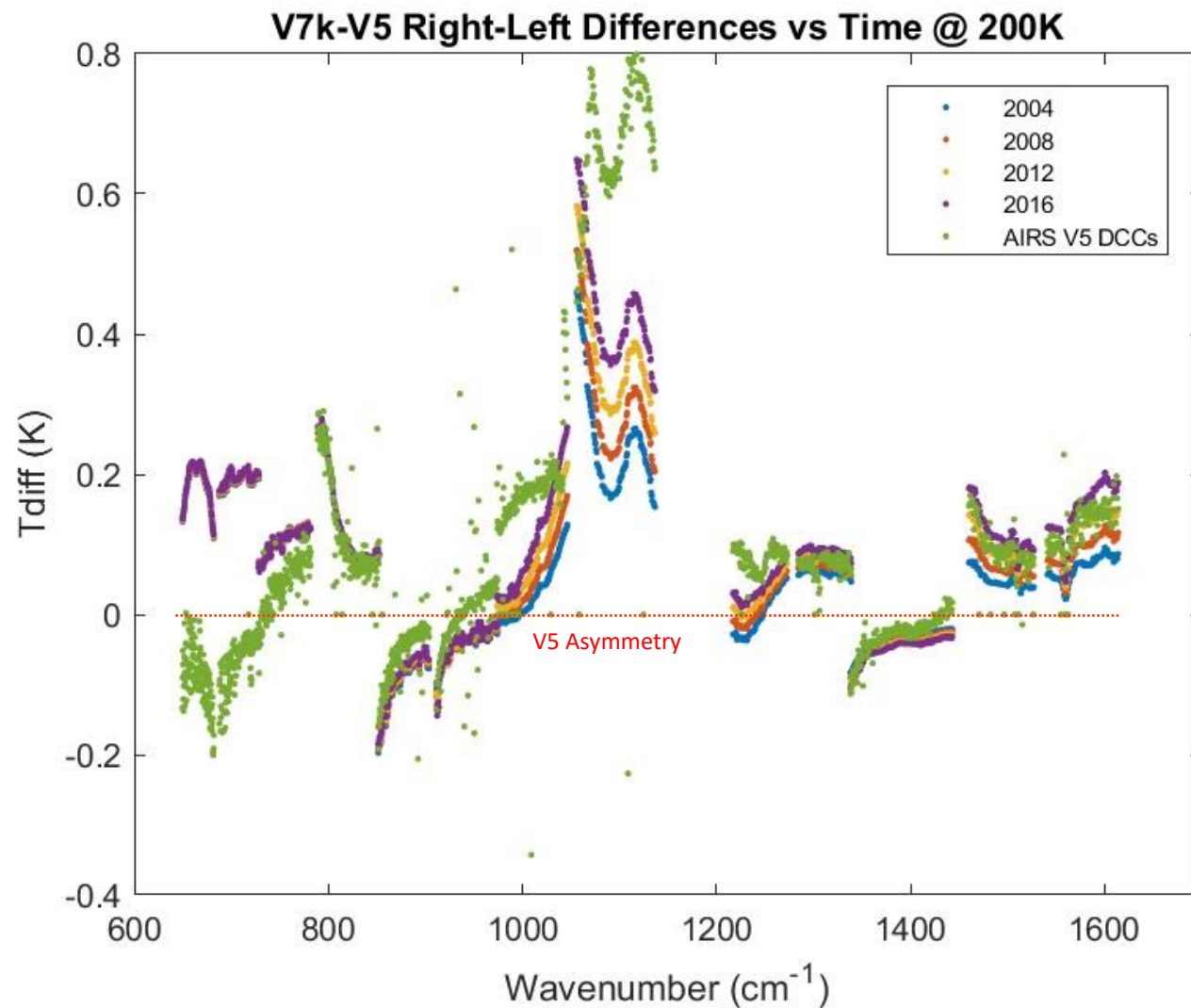
Polarization Parameters Calculated Over Mission Life



AIRS Polarization Product(prpt) for Module Center Channels

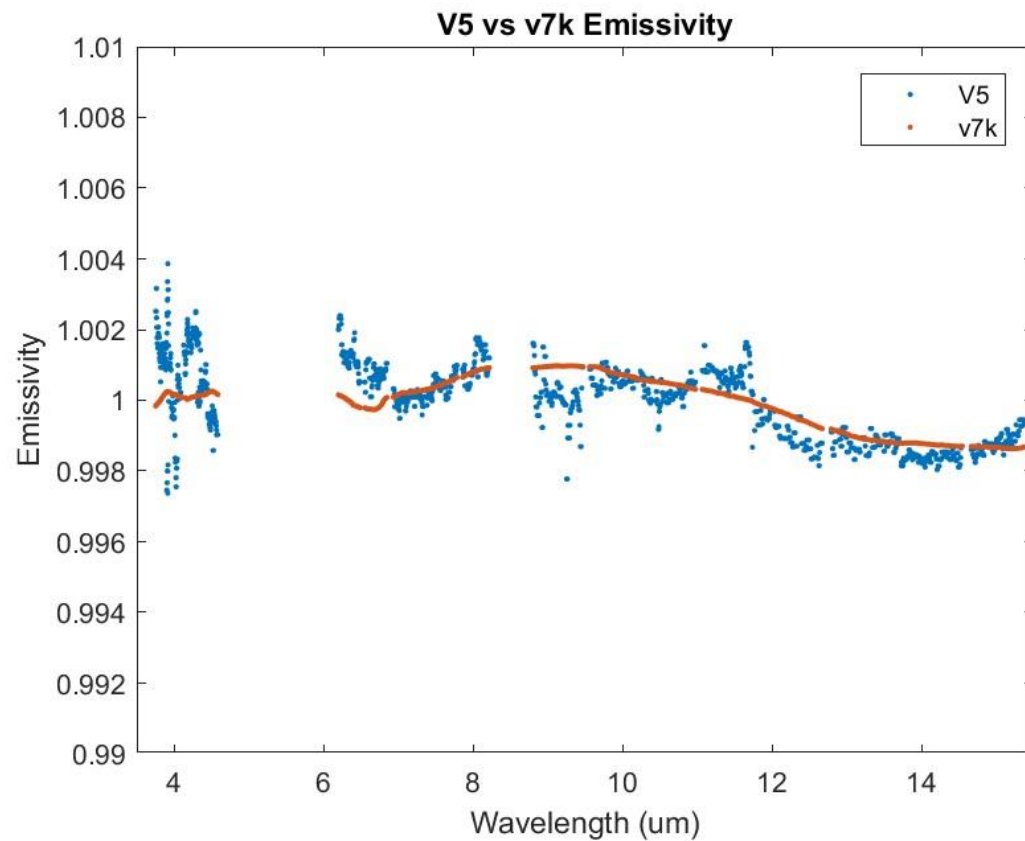


Polarization and Phase in V7k are Time Dependent Using Linear Fit Offset and Slope

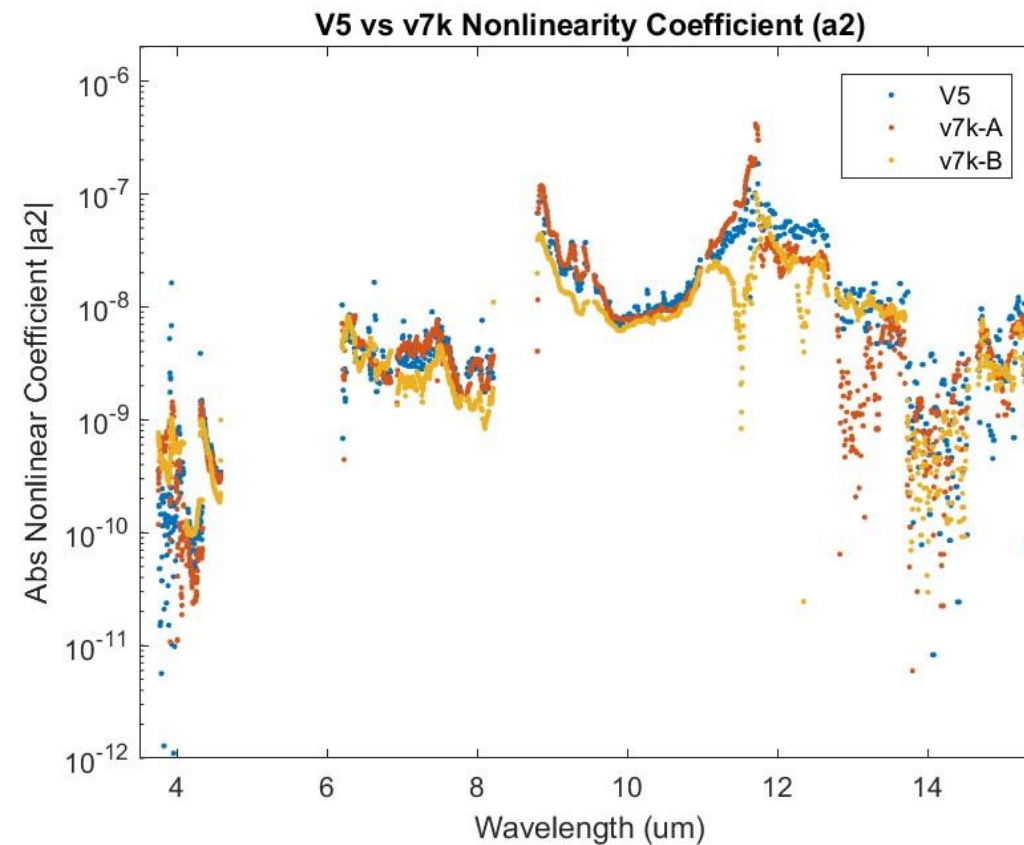


Changes Made to Emissivity and Nonlinearity

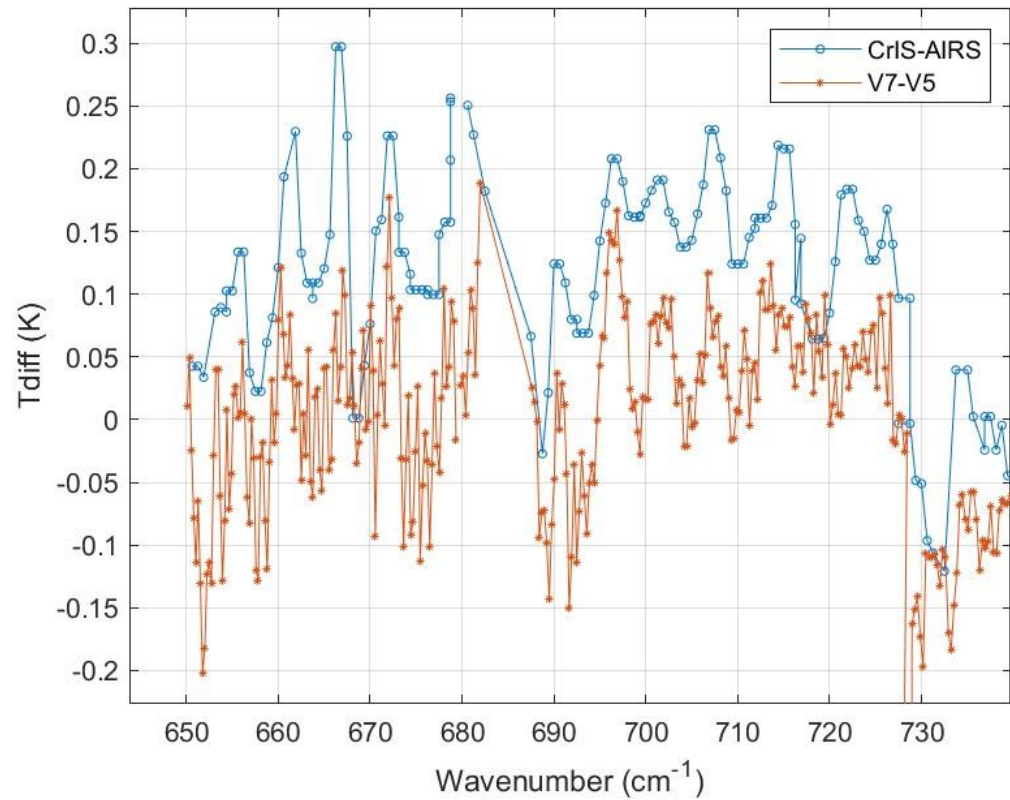
V7k Emissivity Smoothed over Channels



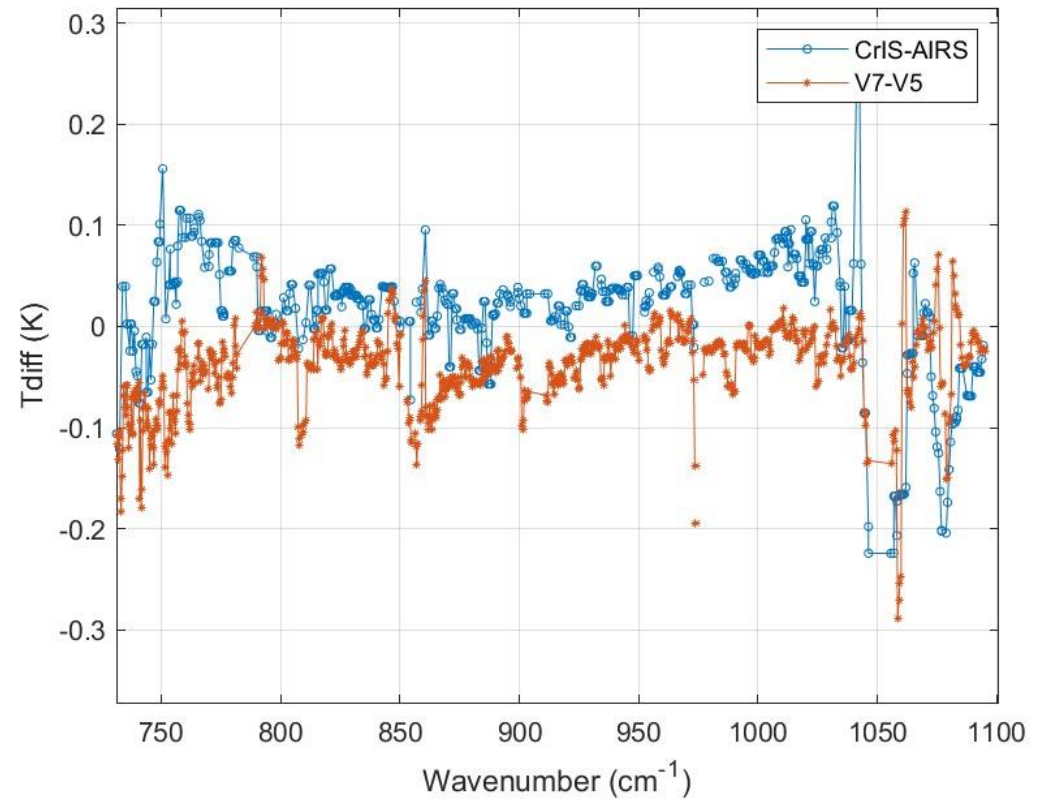
V7k Uses Separate A & B Nonlinearity



VLWIR



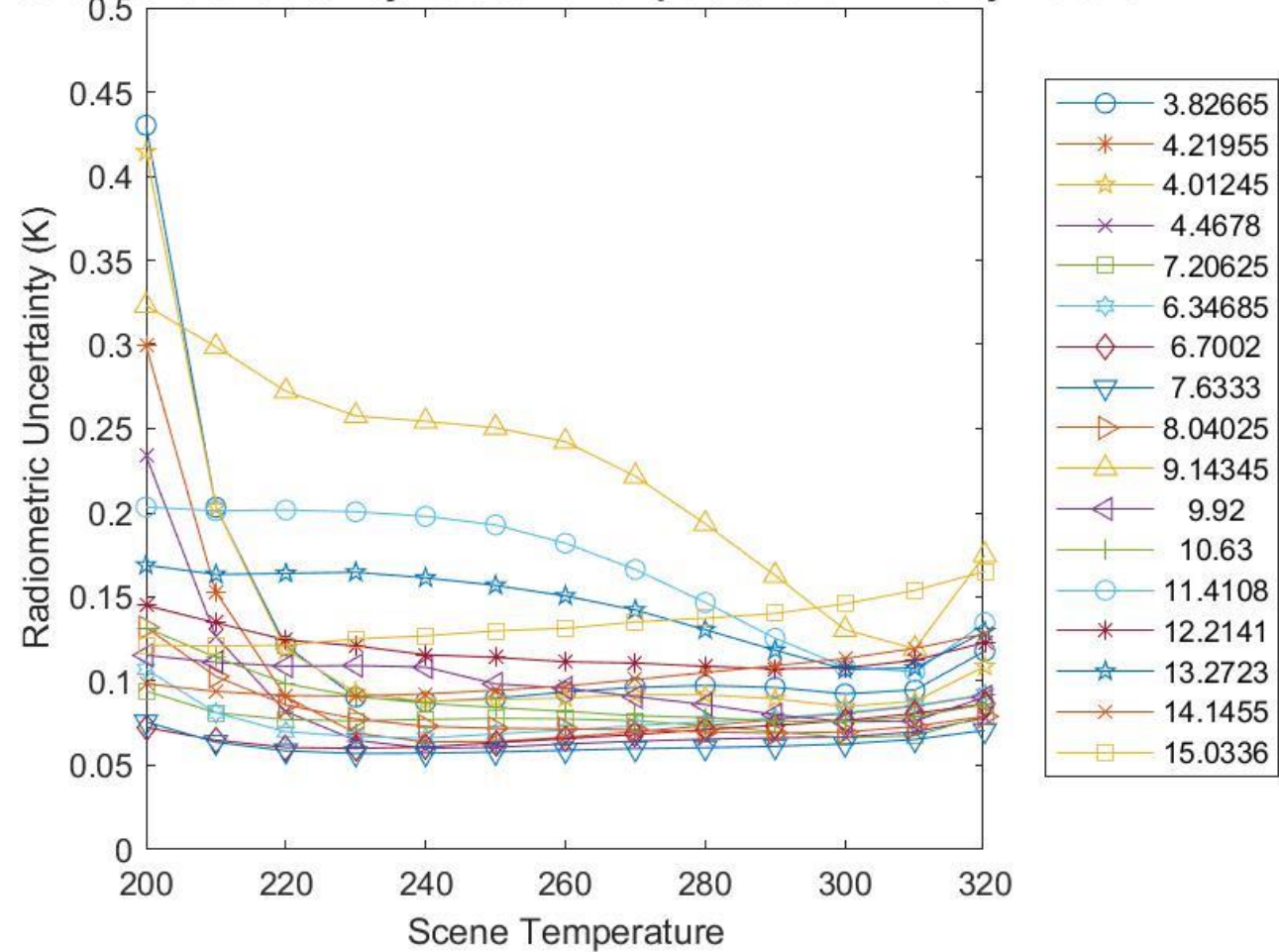
LWIR



Radiometric Uncertainty Inputs at 9.14 μm
and Impact to Calibration at 260K

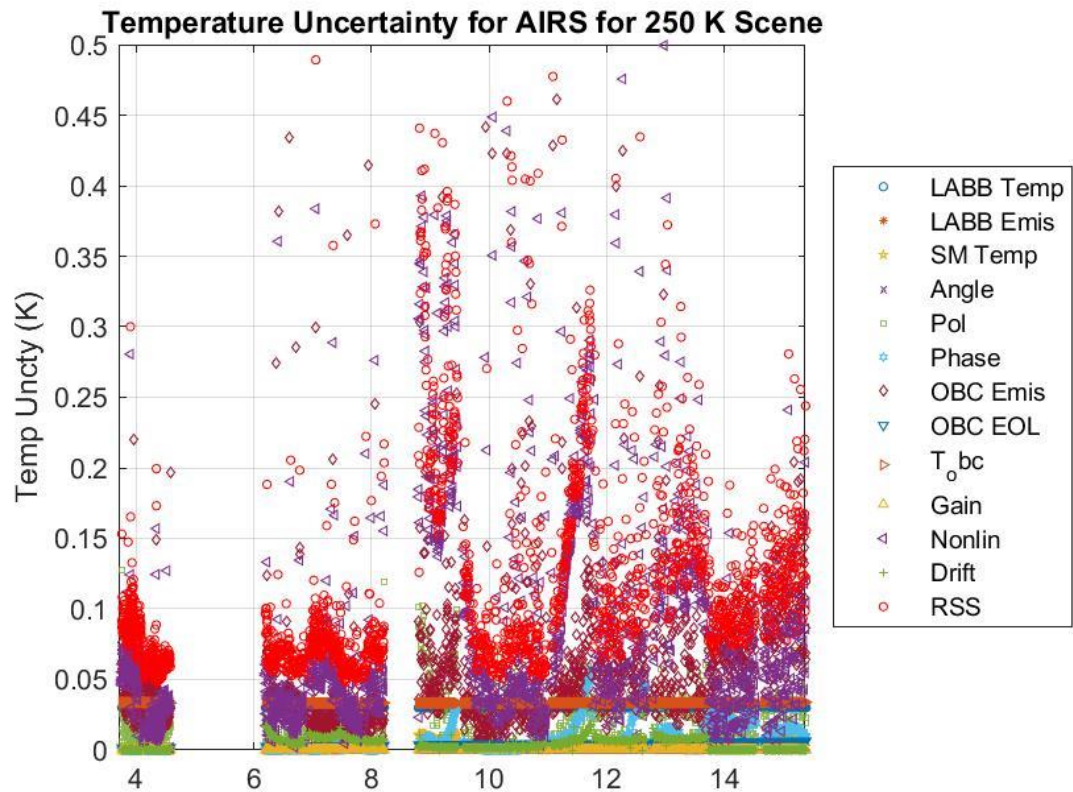
Parameter	1- σ Parameter Uncity	1- σ Radiometric Uncity (K)
Version	V5	V5
Uncertainty in LABB Temperature	0.03K	0.03K
Uncertainty in LABB Emissivity	0.00005	0.002K
Uncertainty in Scan Mirror Temperature	1.25K	0.006K
Uncertainty in Polarization Amplitude	.0009	0.04K
Uncertainty in Polarization Phase	.08	0.005K
Uncertainty in OBC Blackbody Emissivity	.002	0.07K
Uncertainty in OBC Blackbody Emissivity (EOL)	0.0001	0.004K
Uncertainty in OBC Blackbody Temperature	0.05K	.04K
Uncertainty in Nonlinearity	2.7%*	0.21K
Uncertainty in drift in space view	0.04dn	0.001K
Total Uncertainty at 260K		0.24K

Radiometric Uncertainty vs Scene Temperature for AIRS by Module

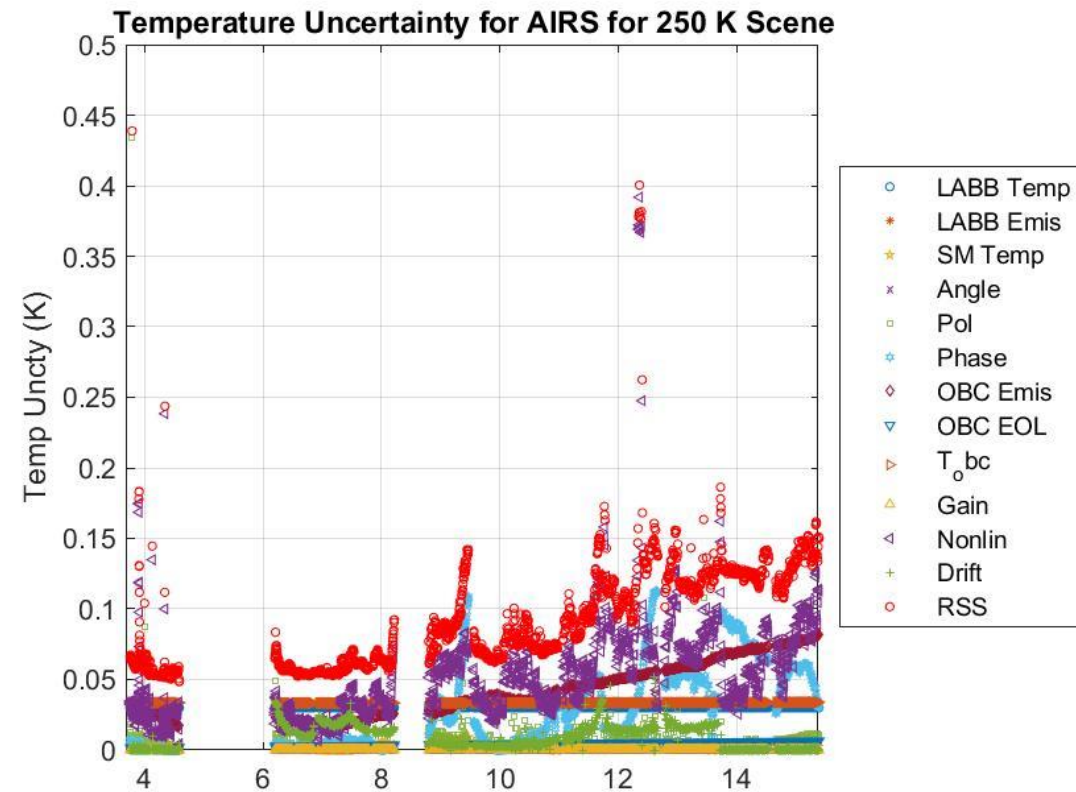


V7k Reduces Radiometric Uncertainty (1σ)

V5



V7k



Version 6.0

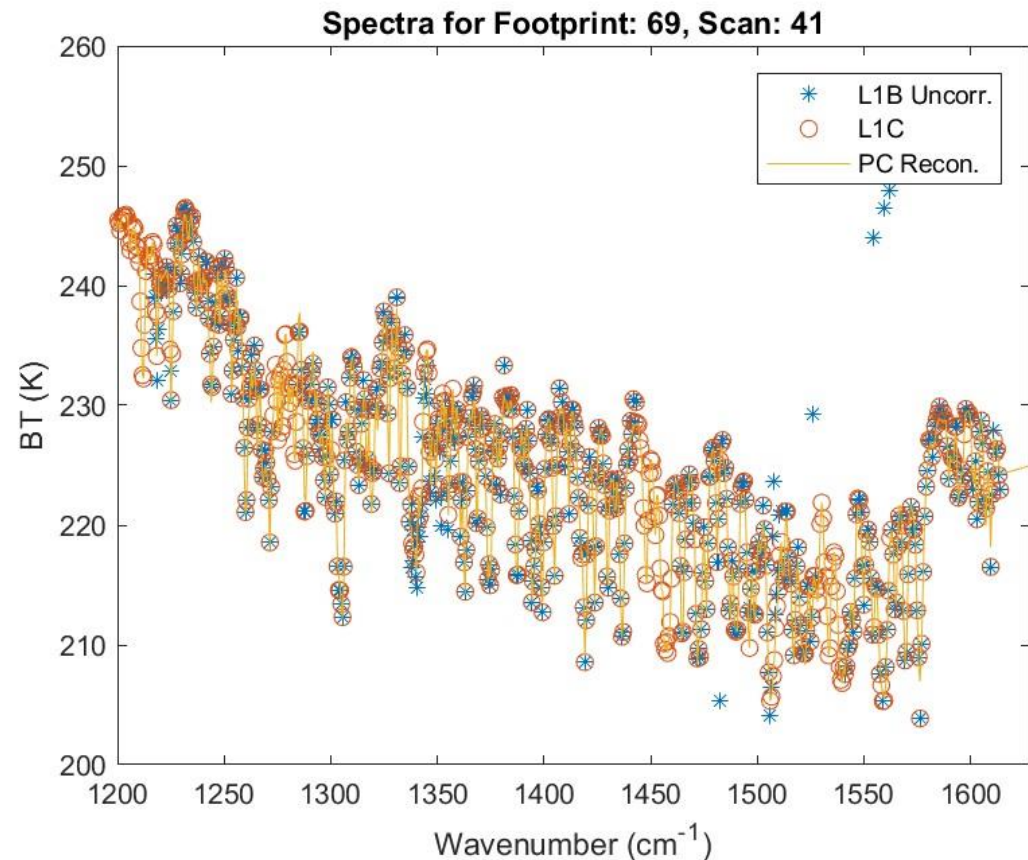
- Designed to facilitate use of AIRS Level 1 radiances
 - For Comparison to other Hyperspectral IR
 - For Comparison to Broadband Imagers
 - For Ingest by L2 Retrievals
- Version 5 L1B for all “good” channels
- Fills Dead Channels with PC Reconstruction (PCR)
- Fills bad Cij (Co-registration) Pixels with PCR
- Fills Gap with PCR
- Fills Very High Noise Pixels with PCR
- V6.0 L1C is available only on a limited basis

Version 6.6

- Same as Version 6.0 but with...
- Constant Frequency Grid (does not change with time)
- V6.6 L1C will be used to support LLS’s CHIRP product

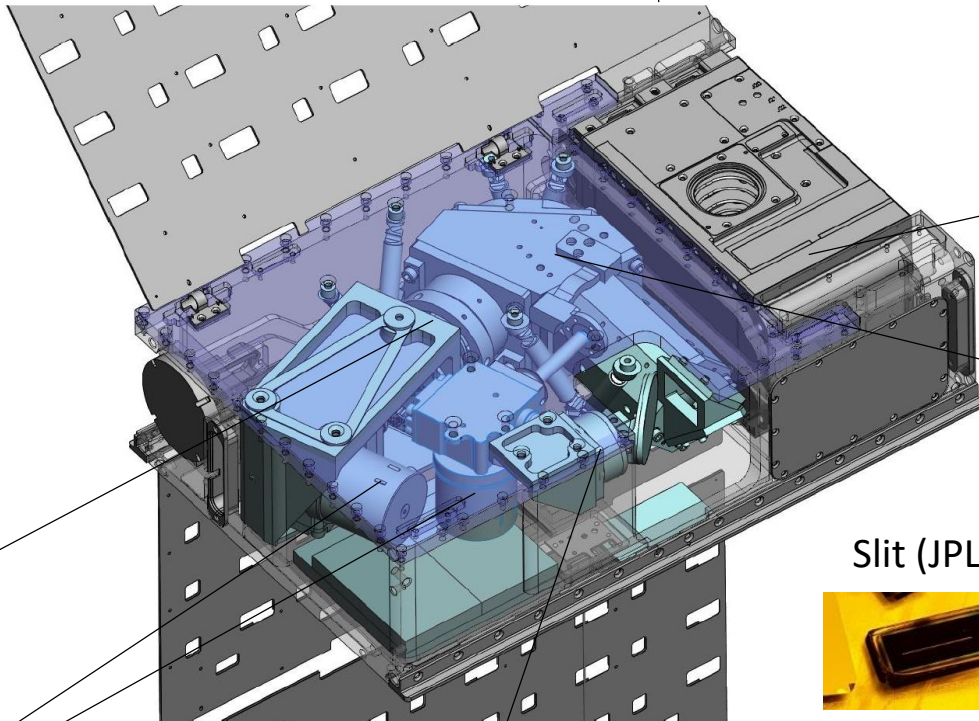
Version 7.0

- Same as Version 6.6 but with
- Updated filling algorithms.
- Version 7k Radiometric and Polarization Coefficients
- netCDF Output

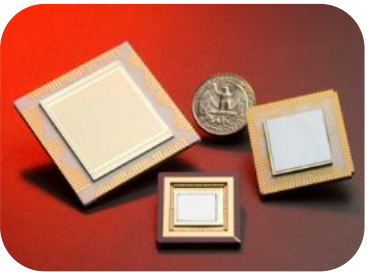


AIRS.2014.03.01.124.L1C.AIRS_Rad.v6.1.0.2.X14295151145.hdf

CubeSat IR Atmospheric Sounder (CIRAS) at JPL



FPA
HOTBIRD
(JPL)



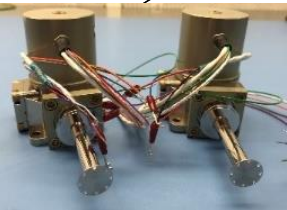
Camera Electronics
(IR Cameras)



Dewar (IDCA)
(IR Cameras)



Cryocoolers +
Electronics
(Ricor K508N)



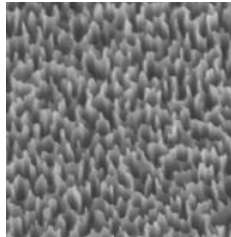
Payload
Electronics



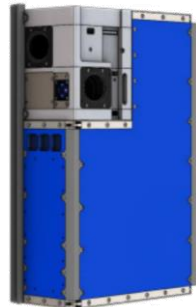
Stepper Motor +
Mirror
(Lin Eng)



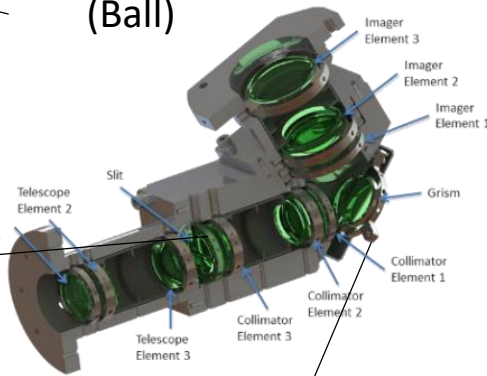
Blackbody
Assembly
Black Silicon



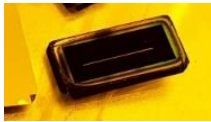
Spacecraft
(BCT)



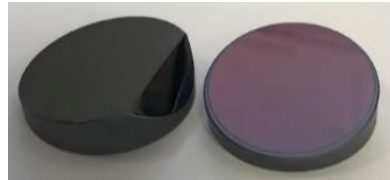
Optics Assembly
(Ball)



Slit (JPL)



Immersion Grating
(JPL)



Hardware Available

Designs Available

- AIRS on Aqua Healthy. Expect mission to continue beyond 2022.
- Continuity from operational weather sounders: CrIS and IASI
- AIRS designed and tested to produce SI traceable radiances
- Radiometric Calibration Coefficients Updated
 - Current operational version using pre-launch coefficients (V5)
 - Updates provided to polarization, emissivity, nonlinearity (V7k)
- V7k Improvements seen in:
 - Reduced Bias and L/R Assymetry in Cold Scenes
 - Agrees better with CrIS
 - Reduction in Radiometric Accuracy
- AIRS Level 1C Fills Bad Channels and Gaps in Spectra. Easier than L1B
- New Technologies enable IR sounding in CubeSat with full aperture cal